

A little geometry

Frame of reference is required:

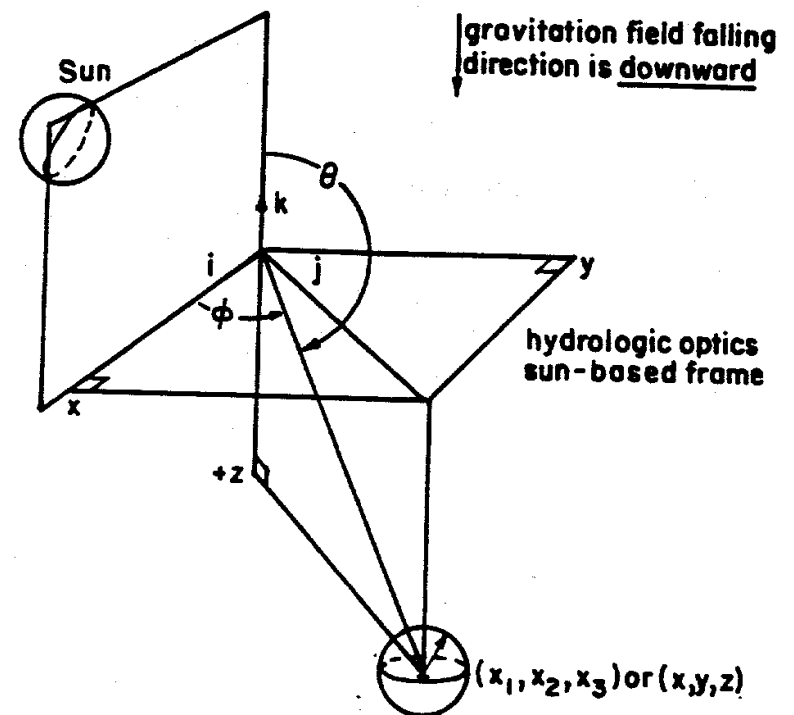
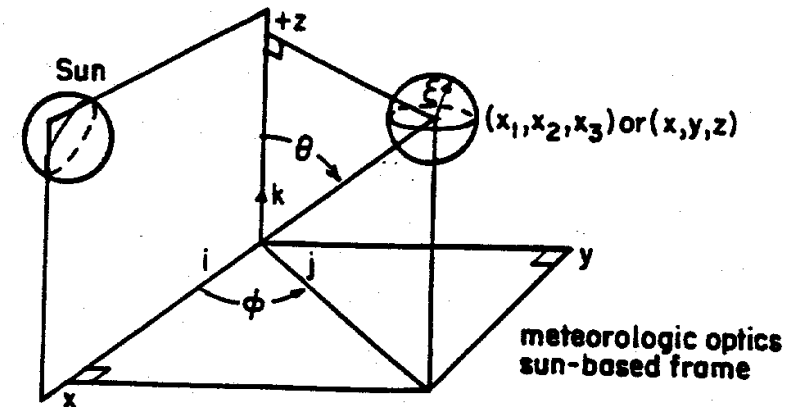
We assume a cartesian coordinate system and orientate this such that the x-axis points to the sun

We will encounter two forms of this 'sun-based' frame of reference - one that assumes the z-axis to be altitude (I.e increasing upwards) and another with z increasing downwards (and thus a measure of depth).

We can specify any position in this space as given by a general position vector \vec{r} by the coordinates of its tip (x,y,z)

The unit vectors $\hat{i}, \hat{j}, \hat{k}$ are the unit vectors defined along the x,y, and z axes

Two angles are also elementary to our considerations - the zenith angle θ and azimuth angle ϕ (measured from the x axis)



Direction

Consider a hypothetical unit sphere about some origin. We define the direction vector as that unit vector that extends from the origin to some point (x,y,z) on that sphere.

$$\hat{\xi} = \frac{\vec{r}}{|\vec{r}|} = \frac{(x, y, z)}{(x^2 + y^2 + z^2)^{1/2}}$$

For a vector of unit length ($|\vec{r}| = 1$):

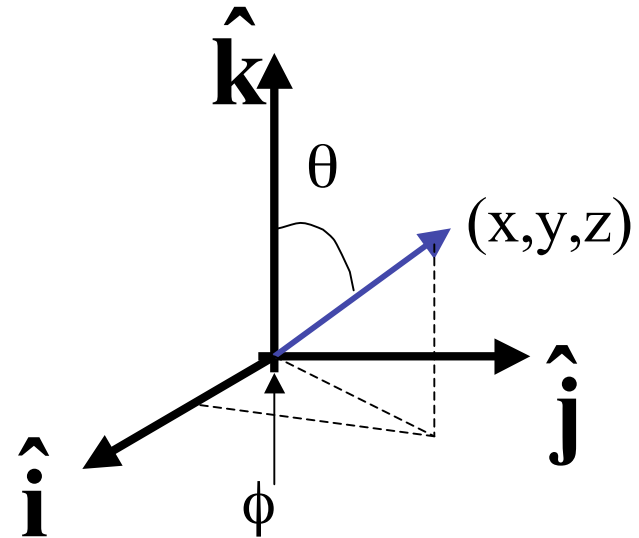
$$x = \vec{r} \cdot \hat{i} = \cos \phi \sin \theta$$

$$y = \vec{r} \cdot \hat{j} = \sin \phi \sin \theta$$

$$z = \vec{r} \cdot \hat{k} = \cos \theta = \mu$$

So that we can express a unit direction vector

$$\hat{\xi} = (\cos \phi \sin \theta, \sin \phi \sin \theta, \cos \theta)$$



Important application - angle formed between two directions

$$\cos \Theta = \hat{\xi} \cdot \hat{\xi}'$$

$$\cos \Theta = \mu \mu' +$$

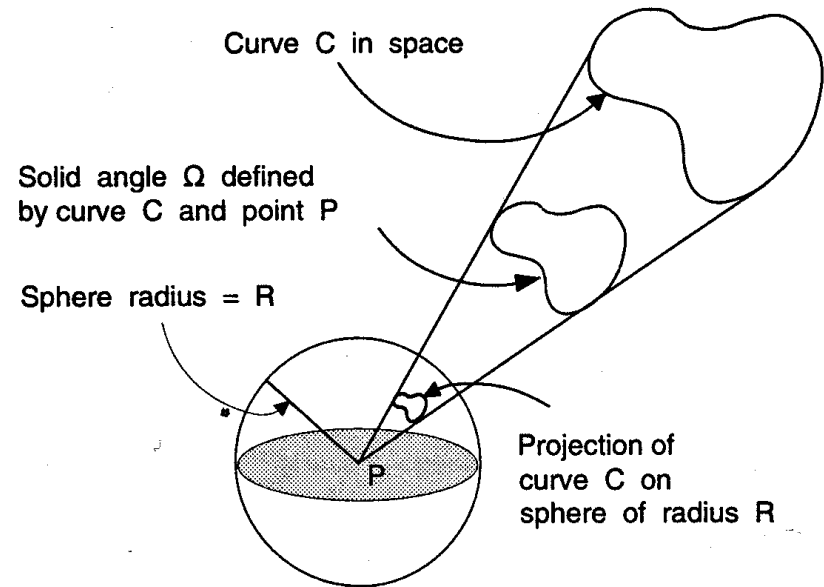
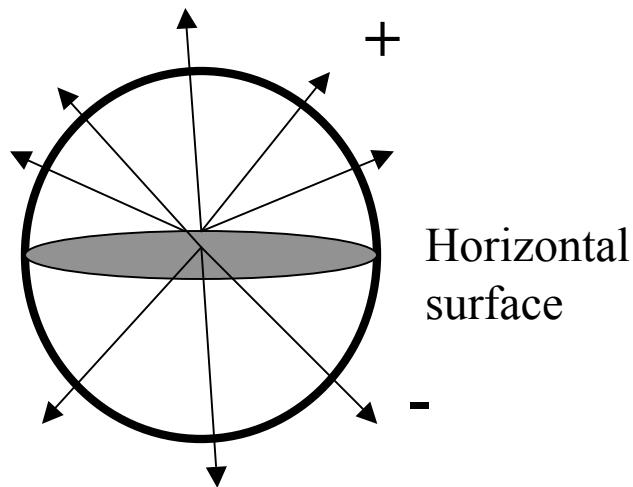
$$(1 - \mu^2)^{1/2} (1 - \mu'^2)^{1/2} \cos(\phi' - \phi)$$

Solid Angle

Consider a unit sphere with some opening of area dA as shown:

solid angle subtended by the opening at origin = area of the opening on the unit sphere

Small element of area is $da \times db = \sin\theta d\theta \times d\phi = d\Omega$



Establish this unit sphere relative to some direction, such as local vertical - establishes two hemispheres (upper and lower)

Solid Angle Examples

Example 2.2: Solid Angle

- The solid angle of a spherical cap defined by the angle θ is

$$\begin{aligned}\Omega(D) &= \int_0^\theta \sin \theta d\theta \int_0^{2\pi} d\phi \\ &= 2\pi[1 - \cos \theta]\end{aligned}$$

For small θ , $\cos \theta \rightarrow 1 - \theta^2/2 + \dots$ and

$$\Omega(D) = \pi\theta^2$$

For $\theta = \pi$, the solid angle of a sphere is

$$\Omega(D) = 4\pi.$$

- The solid angle of a spherical segment is

$$\begin{aligned}\Omega(D) &= \int_{\theta_1}^{\theta_2} \sin \theta d\theta \int_0^{2\pi} d\phi \\ &= 2\pi[\cos \theta_1 - \cos \theta_2]\end{aligned}$$

- The solid angle of the sun is

$$\Omega_{\odot} = \pi\theta^2$$

where as we shall see later, $\theta \approx r_{\odot}/R_{SE}$, and

$$\Omega_{\odot} = \pi\left(\frac{0.7 \times 10^6}{1.5 \times 10^8}\right)^2 \approx 0.684 \times 10^{-4} \text{ steradian}$$

Radiant flux - a basic quantity*

Radiant flux along \vec{r}

$$P(\vec{r}) = h\nu \times n(\vec{r}) \times c \times dA$$

$W(\mu\text{m})^{-1}$

Photon energy Photon # density Photon speed Area element orthogonal to \vec{r}

$$F = P/dA = n(r)c h\nu$$

$$n(r)c = F/h\nu$$

For $F = 0.1 \text{ Wm}^{-2} (\mu\text{m})^{-1}$, $\lambda = 0.5 \mu\text{m}$, $c = 3 \times 10^8 \text{ ms}^{-1}$, $h = 6.6 \times 10^{-34} \text{ Js/photon}$, then

$$n(r)c = 0.1 \times 0.5 \times 10^{-6} / (6.6 \times 10^{-34} \times 3 \times 10^8) = 2.5 \times 10^{17}$$

photons of $\lambda = 0.5 \mu\text{m}$ flow per second through a unit area producing 0.1

Watt of power per $(\mu\text{m})^{-1}$

* but one not directly measurable

Basic measurement concepts -radiance

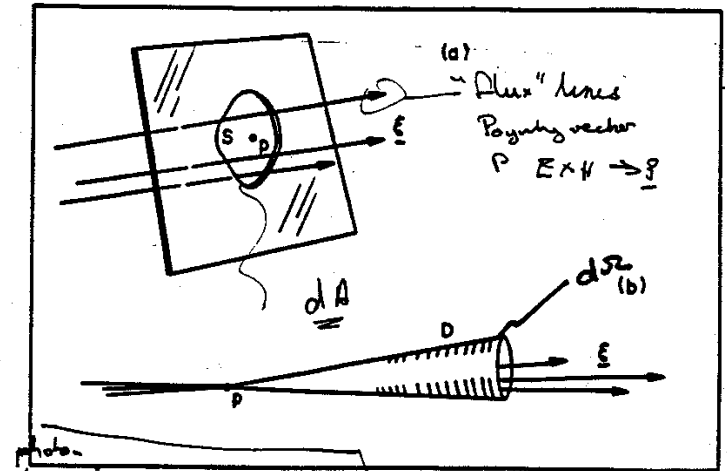
Two quantities thus follow from $P(\vec{r})$:

Area density radiant flux:

$$F_A = P(\vec{r})/dA (\xi) \quad \text{Wm}^{-2} (\mu\text{m})^{-1}$$

Solid angle density

$$F_\Omega = P(\vec{r}) / d\Omega (\vec{\xi}) \quad \text{Wster}^{-1} (\mu\text{m})^{-1}$$



Basic quantity measured is the **radiance**

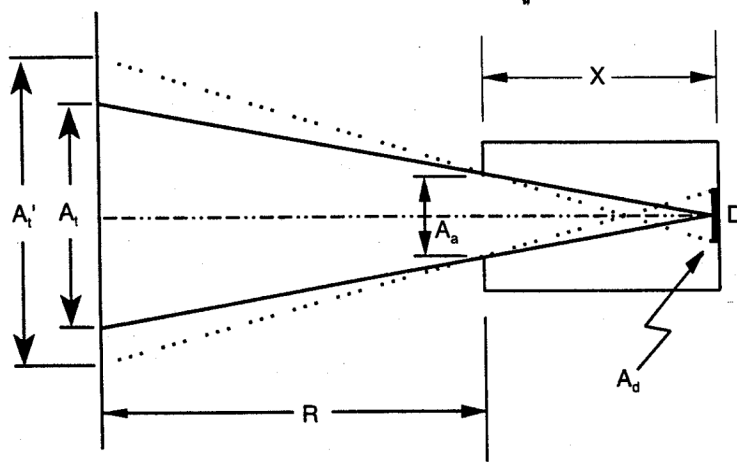
$$I = P/T \quad \text{Wm}^{-2}\text{ster}^{-1} (\mu\text{m})^{-1}$$

where $T = dA \times d\Omega$

is the instrument throughput.

Radiance or intensity is fundamental since we can measure it and all other relevant parameters of interest to us derive from it.

Basic measurement concepts -radiometer



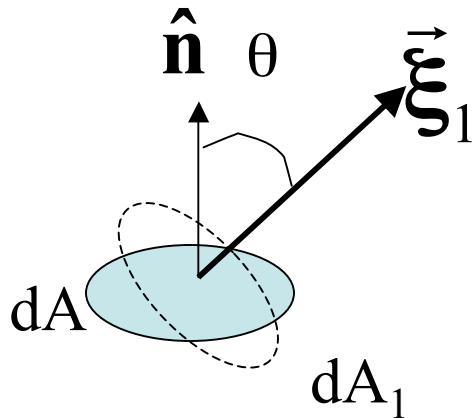
Key point: radiance is a 'field' quantity being independent of the distance between the instrument and source assuming the FOV is uniformly filled.

Problem: a simple radiometer is pointed at a wall as shown. The radiant flux received by a detector D of area A_d at the base of a black tube of length X and aperture area A_a as shown is P . Assuming $A_d \ll X^2$ and $A_a \ll R^2$ what is the averaged radiance of the wall?

Solution: The solid angle subtended by the entrance aperture at the center of the detector is A_a/X^2
The radiance is $P/(A_d A_a/X^2)$

Note: (i) the area of the field of view (target area) is A_t at the point p. The total area as seen by the whole detector is A_t' which is larger than A_t . (ii) The radiance is independent of R .

From radiance to irradiance (flux)



- The radiance along direction $\vec{\xi}_1$ is

$$I_1 = \frac{P_1}{dA_1 d\Omega}$$

- The projection of dA_1 onto the surface perpendicular to $\hat{\mathbf{n}}$ is

$$dA = dA_1 \hat{\mathbf{n}} \cdot \vec{\xi}_1$$

- The total energy through dA per unit area

$$F = \frac{P_1}{dA} = I_1 (\hat{\mathbf{n}} \cdot \vec{\xi}_1) d\Omega(\vec{\xi}_1) \quad \text{Wm}^{-2} (\mu\text{m})^{-1}$$

We call F the flux or irradiance

For n sources of radiance $I_j, j=1\dots n$ along the n directions $\vec{\xi}_j, j=1\dots n$ illuminating the surface, the total rate of energy flow per unit area through the surface dA is just the superposition of each individual source

From radiance to irradiance (flux)

For n sources $\mathbf{I}_1 = \mathbf{I}(\vec{\xi}_1)$

$$\mathbf{F}(\hat{\mathbf{n}}) = \frac{\mathbf{P}(\hat{\mathbf{n}})}{dA} = \mathbf{I}_1 \hat{\mathbf{n}} \cdot \vec{\xi}_1 d\Omega(\vec{\xi}_1) + \mathbf{I}_2 \hat{\mathbf{n}} \cdot \vec{\xi}_2 d\Omega(\vec{\xi}_2) + \dots$$

$$\mathbf{F}(\hat{\mathbf{n}}) = \int \mathbf{I}(\vec{\xi}') \hat{\mathbf{n}} \cdot \vec{\xi}' d\Omega(\vec{\xi}')$$

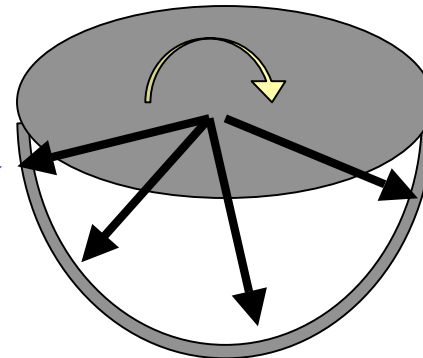
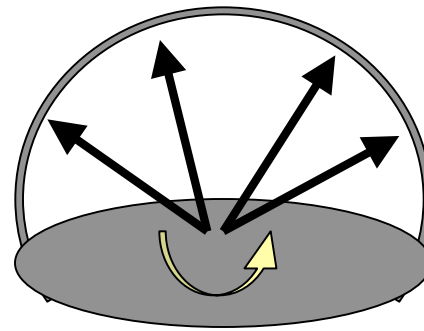
Consider now the special case
with $\hat{\mathbf{n}} = \hat{\mathbf{k}}$

$$d\Omega = \sin \theta d\theta d\phi$$

$$\mathbf{F}^+ = \int_0^{2\pi} d\phi \int_0^{\pi/2} \mathbf{I}(\theta, \phi) \cos \theta \sin \theta d\theta$$

$$(\hat{\mathbf{k}} \cdot \vec{\xi} = \cos \theta)$$

$$\mathbf{F}^- = \int_0^{2\pi} d\phi \int_{-\pi/2}^0 \mathbf{I}(\theta, \phi) \cos \theta \sin \theta d\theta$$



Scattering phase function

Consider an experiment that places a detector at some distance from a scattering volume. We seek to measure the scattered intensity at all such points located around an imaginary sphere that surrounds the scatterer

As before:

$$E_{inc} = E_0 e^{-ikz + i\omega t}$$

$$E_{sca} = S(\theta) \frac{e^{-ikr + i\omega t}}{kr}$$

$$E_{sca} = S(\theta) \frac{e^{-ikr + ikz}}{kr} E_0$$

$$I_{sca} = \frac{|S(\theta)|^2}{k^2 r^2} I_0$$

$$I_{sca}(\theta) = \frac{C_{sca} k^2}{4\pi} P(\theta) I_0$$

power received by detector

$$dW = I_{sca} dA = I_{sca} r^2 d\Omega = \frac{|S(\theta)|^2}{k^2} I_0 d\Omega$$

total power scattered

$$W = \int_{\Xi} \frac{|S(\theta)|^2}{k^2} I_0 d\Omega = I_0 C_{sca}$$

thus

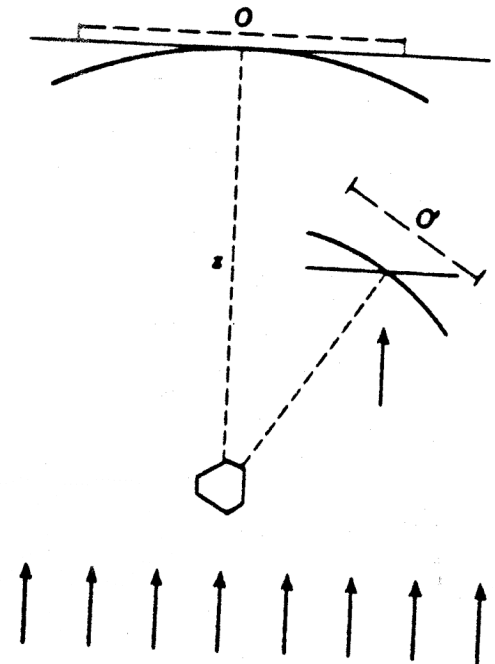
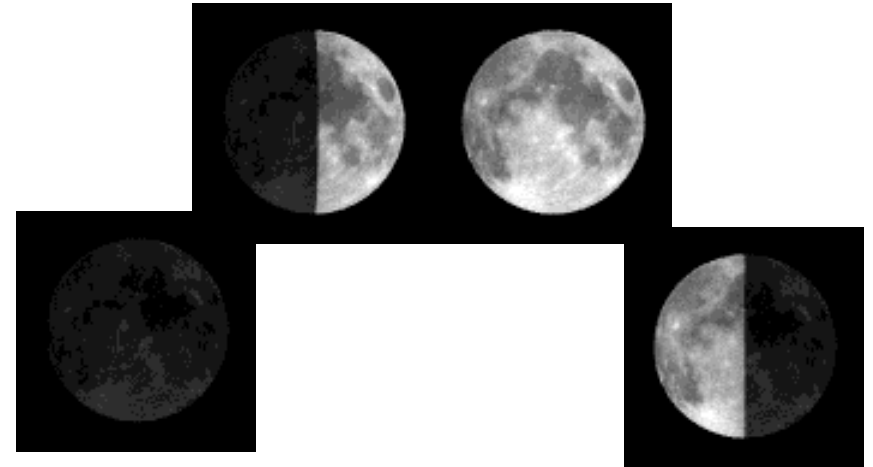
$$C_{sca} = \int_{\Xi} \frac{|S(\theta)|^2}{k^2} d\Omega$$

define

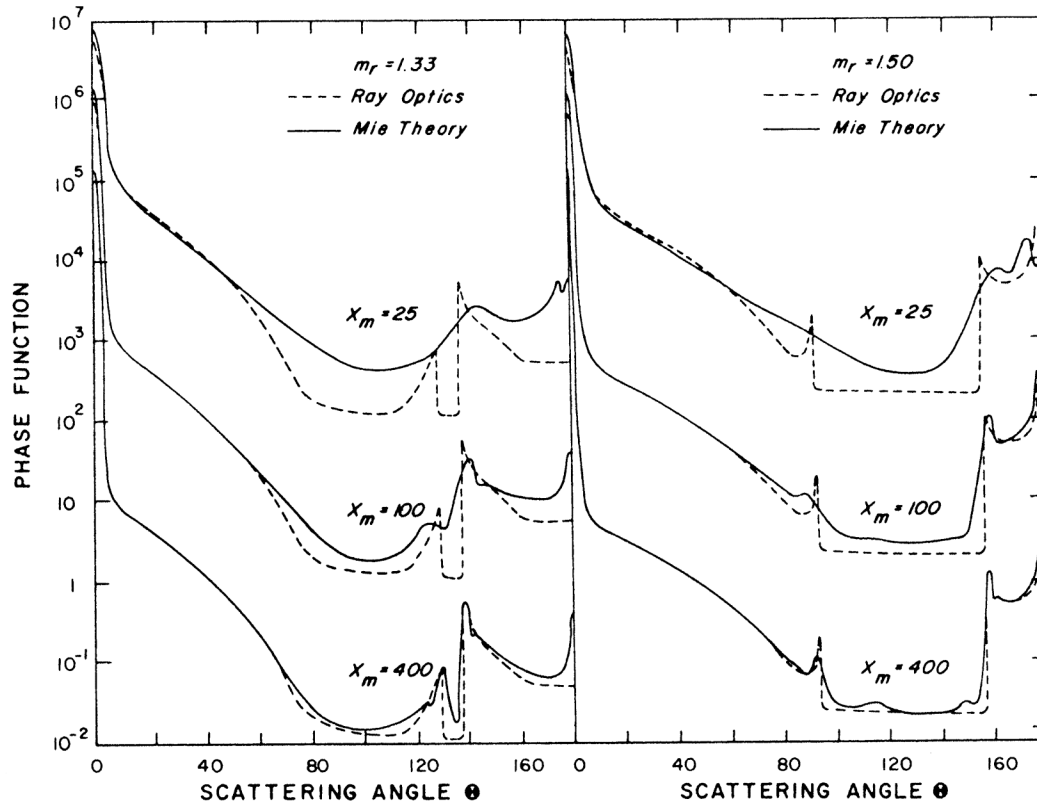
$$\frac{1}{4\pi} P(\theta) = \frac{|S(\theta)|^2}{C_{sca} k^2}$$

by definition

$$\frac{1}{4\pi} \int_{\Xi} P(\theta) d\Omega = 1$$



Examples of phase function



Points to note: the extent of forward scattering & how it increase with x

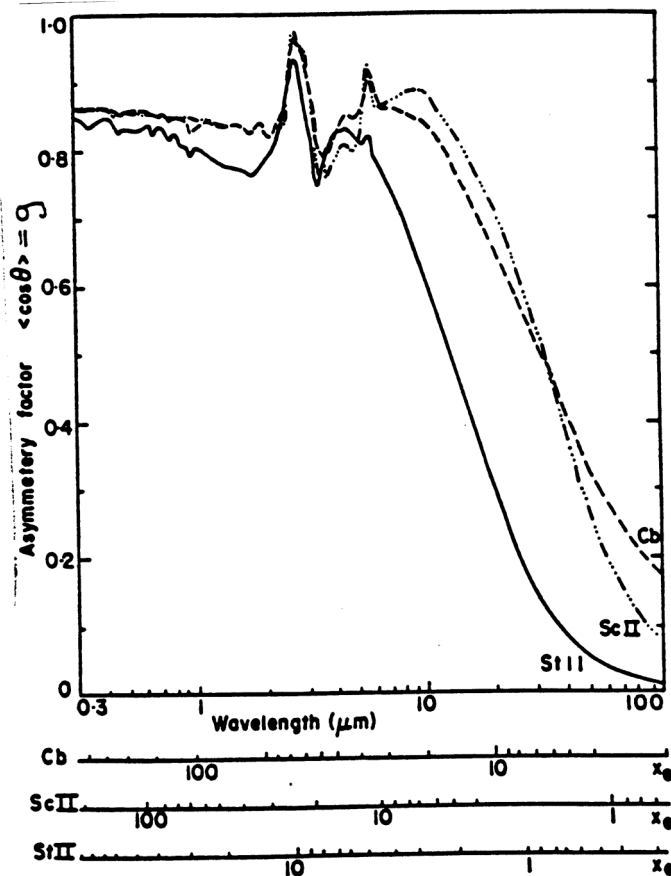
Optical phenomena like rainbow and glory and where they appear

Smoothing of scattering function for polydispersions

Anticipate the effects of particle absorption

<http://members.tripod.com/~regenbogen/indexe.htm>

Properties of the phase function



$$g = \frac{1}{2} \int_{-1}^{+1} P(\cos \Theta) \cos \Theta d \cos \Theta$$

asymmetry parameter

$g=1$ pure forward scatter

$g=0$ isotropic or symmetric

(e.g

Rayleigh)

$g=-1$ pure backscatter

Other properties of the phase function

Define forward scatter & backscatter as

$$f = \frac{1}{4\pi} \int_0^{2\pi} \int_0^1 P(\theta, \phi) d\cos\theta = \frac{1}{2} P(0)$$

$$b = \frac{1}{4\pi} \int_0^{2\pi} \int_{-1}^0 P(\theta, \phi) d\cos\theta = \frac{1}{2} P(180)$$

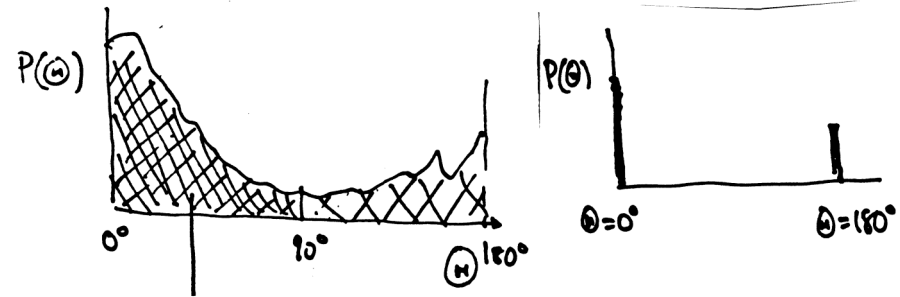
$f + b = 1$ (normalization condition)

For a simple delta phase function

$$g = \frac{1}{4\pi} \int_0^{2\pi} \int_{-1}^1 P(\theta, \phi) \cos\theta d\cos\theta = f - b$$

$$b = \frac{1}{2}(1 - g)$$

$$f = \frac{1}{2}(1 + g)$$



Other (mathematical) forms of phase function

$$P(\cos\Theta) = \sum_{\ell=0}^N \chi_{\ell} P_{\ell}(\cos\Theta)$$

a particular form of function has

$$\chi_{\ell} = \frac{(2\ell + 1)}{2} g^{\ell}$$

and with $N = 1$

$$P(\cos\Theta) = \left(1 + \frac{3}{2}g \cos\Theta\right)$$

Legendre polynomial expansion

Legendre Polynomial of degree ℓ